The Schwarzschild solution and its implications for gravitational waves

By Stephen J. Crothers

"Schwarzschild's Solution" is not Schwarzschild's solution

$$ds^{2} = \left(1 - \frac{2m}{r}\right)dt^{2} - \left(1 - \frac{2m}{r}\right)^{-1}dr^{2} - r^{2}\left(d\theta^{2} + \sin\theta^{2}d\varphi^{2}\right)$$

David Hilbert's corruption of Schwarzschild's solution, wherein it is claimed, without proof, that r can go down to zero, producing an event horizon at r = 2m and an infinitely dense point-mass singularity at r = 0.

Schwarzschild's Solution

$$ds^{2} = \left(1 - \frac{\alpha}{R}\right) dt^{2} - \left(1 - \frac{\alpha}{R}\right)^{-1} dR^{2} - R^{2} \left(d\theta^{2} + \sin\theta^{2} d\varphi^{2}\right)$$

$$R = R(r) = (r^3 + \alpha^3)^{\frac{1}{3}}, \qquad 0 < r < \infty.$$

There is no black hole.

3-D Spherically Symmetric Metric Manifolds (Euclidean Space)

$$ds^{2} = dr^{2} + r^{2} \left(d\theta^{2} + \sin \theta^{2} d\varphi^{2} \right)$$
$$0 \le r < \infty.$$

$$R_p = \int_0^r dr = r,$$
 $A_p = 4\pi r^2,$ $V_p = \frac{4}{3}\pi r^3$

Non-Euclidean Metric (Spherically Symmetric)

$$ds^{2} = dR_{p}^{2} + R_{c}^{2} \left(d\theta^{2} + \sin^{2} \theta d\varphi^{2} \right)$$
$$= \Psi(R_{c}) dR_{c}^{2} + R_{c}^{2} \left(d\theta^{2} + \sin^{2} \theta d\varphi^{2} \right),$$

$$R_c = R_c(r), \qquad R_c(0) \le R_c(r) < \infty.$$

 $R_c(r)$ and $\Psi(R_c)$ are *a priori* unknown functions.

Non-Euclidean Geometric Relations

$$R_{p} = \int_{0}^{R_{p}} dR_{p} = \int_{R_{c}(0)}^{R_{c}(r)} \sqrt{\Psi(R_{c}(r))} dR_{c}(r) = \int_{0}^{r} \sqrt{\Psi(R_{c}(r))} \frac{dR_{c}(r)}{dr} dr,$$

$$C_{p} = R_{c}(r) \int_{0}^{2\pi} d\varphi = 2\pi R_{c}(r), \qquad A_{p} = R_{c}^{2}(r) \int_{0}^{\pi} \sin \theta \, d\theta \int_{0}^{2\pi} d\varphi = 4\pi R_{c}^{2}(r),$$

$$V_{p} = \int_{0}^{R_{p}} R_{c}^{2}(r) dR_{p} \int_{0}^{\pi} \sin \theta \, d\theta \int_{0}^{2\pi} d\varphi = 4\pi \int_{R_{c}(0)}^{R_{c}(r)} \overline{\Psi(R_{c}(r))} R_{c}^{2}(r) dR_{c}(r)$$
$$= 4\pi \int_{0}^{r} \sqrt{\Psi(R_{c}(r))} R_{c}^{2}(r) \frac{dR_{c}(r)}{dr} dr$$

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Spherically Symmetric Surfaces

$$ds^{2} = r^{2} \left(d\theta^{2} + \sin^{2} \theta d\varphi^{2} \right)$$

What is *r* ?

$$ds^{2} = R_{c}^{2} \left(d\theta^{2} + \sin^{2} \theta d\varphi^{2} \right)$$
$$R_{c} = R_{c} (r),$$
What is $R_{c}(r)$?

Gaussian Curvature (K) $K = \frac{R_{1212}}{K_{1212}}$ $R_{\mu\nu\rho\sigma} = g_{\mu\gamma}R_{.\nu\rho\sigma}^{\gamma} \qquad \qquad R_{.212}^{1} = \frac{\partial\Gamma_{22}^{1}}{\partial r^{1}} - \frac{\partial\Gamma_{21}^{1}}{\partial r^{2}} + \Gamma_{22}^{k}\Gamma_{k1}^{1} - \Gamma_{21}^{k}\Gamma_{k2}^{1}$ $\Gamma_{ij}^{i} = \Gamma_{ji}^{i} = \frac{\partial \left(\frac{1}{2} \ln |g_{ii}|\right)}{\sum_{i}^{j}}$ $\Gamma_{jj}^{i} = -\frac{1}{2g_{ii}} \frac{\partial g_{jj}}{\partial x^{i}} \quad (i \neq j)$ i, j, k = 1, 2, 3 $x^1 = \theta, x^2 = \phi$ all other Γ^{i}_{ik} vanish

Gaussian Curvature (K) (Spherically Symmetric Surface)

$$K = \frac{1}{R_c^2}$$

Gaussian curvature is an **intrinsic** geometrical property of a surface, completely independent of any embedding space (**Theorema Egregium** of C. F. Gauss). Gaussian curvature is a bending invariant. Its identity is **not altered** if the surface is embedded in a higher-dimensional space.

Standard Derivation

$$ds^{2} = e^{\lambda} dt^{2} - e^{\beta} dr^{2} - R^{2} \left(d\theta^{2} + \sin^{2} \theta d\varphi^{2} \right)$$

Generalisation of the Minkowski squared arc-length, where R=R(r).

$$ds^{2} = e^{\lambda} dt^{2} - e^{\beta} dr^{2} - r^{2} \left(d\theta^{2} + \sin^{2} \theta d\varphi^{2} \right)$$

Reduced generalisation of the Minkowski squared arc-length, where it is **asserted** that R(r) = r and so that R(0) = 0, **without proof**.

For λ,β real valued functions of *r* alone, $e^{\lambda} > 0$ and $e^{\beta} > 0$ by construction: e^{λ} and e^{β} cannot change sign or become zero. The metric signature is thereby fixed at (+,-,-,-).

General Line-Element for
$$Ric = 0$$

 $ds^2 = A(R_c)dt^2 - B(R_c)dR_c^2 - R_c^2(d\theta^2 + \sin^2\theta d\varphi^2)$
 $A(R_c), B(R_c), R_c(r) > 0$

General Solution for
$$Ric = 0$$

$$ds^{2} = \left(1 - \frac{\alpha}{R_{c}}\right) dt^{2} - \left(1 - \frac{\alpha}{R_{c}}\right)^{-1} dR_{c}^{2} - R_{c}^{2} \left(d\theta^{2} + \sin^{2}\theta d\varphi^{2}\right)$$
$$R_{c} \left(0\right) \le R_{c} \left(r\right) < \infty$$

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Proper Radius

$$R_{p} = \int \frac{dR_{c}}{\sqrt{1 - \frac{\alpha}{R_{c}}}} = \sqrt{R_{c}(R_{c} - \alpha)} + \alpha \ln\left(\sqrt{R_{c}} + \sqrt{R_{c} - \alpha}\right) + k$$

For some r_o , $R_p(r_o) = 0$; hence $R_c(r_o) = \alpha$, $k = -\alpha \ln \sqrt{\alpha}$.

$$R_{p}(r) = \sqrt{R_{c}(R_{c} - \alpha)} + \alpha \ln\left(\frac{\sqrt{R_{c}} + \sqrt{R_{c} - \alpha}}{\sqrt{\alpha}}\right)$$
$$R_{c} = R_{c}(r)$$

Brillouin's Solution
$$ds^{2} = \left(1 - \frac{\alpha}{r + \alpha}\right) dt^{2} - \left(1 - \frac{\alpha}{r + \alpha}\right)^{-1} dr^{2} - (r + \alpha)^{2} \left(d\theta^{2} + \sin^{2}\theta d\varphi^{2}\right)$$
$$0 < r < \infty$$

Droste's Solution

$$ds^{2} = \left(1 - \frac{\alpha}{r}\right) dt^{2} - \left(1 - \frac{\alpha}{r}\right)^{-1} dr^{2} - r^{2} \left(d\theta^{2} + \sin^{2}\theta d\varphi^{2}\right)$$
$$\alpha < r < \infty$$

Admissible form for $R_c(r)$

$$R_{c}(r) = \left(\left| r - r_{o} \right|^{n} + \alpha^{n} \right)^{1/n} = \frac{1}{\sqrt{K(r)}}, \quad r \in \mathbf{R}, \ n \in \mathbf{R}^{+}, \ r \neq r_{o}$$

Solution for $Ric \equiv R_{\mu\nu} = 0$

$$ds^{2} = \left(1 - \frac{\alpha}{R_{c}}\right) dt^{2} - \left(1 - \frac{\alpha}{R_{c}}\right)^{-1} dR_{c}^{2} - R_{c}^{2} \left(d\theta^{2} + \sin^{2}\theta d\varphi^{2}\right)$$

$$R_{c}(r) = \left(\left| r - r_{o} \right|^{n} + \alpha^{n} \right)^{\frac{1}{n}} = \frac{1}{\sqrt{K(r)}},$$

$$r \in \mathbf{R}, n \in \mathbf{R}^+, r \neq r_o$$

r_{o} and *n* entirely arbitrary constants

 $R_c(r_o) = \alpha$ and $R_p(r_o) = 0$ are scalar invariants.

Isotropic Coordinates

$$ds^{2} = \frac{\left(1 - \frac{\alpha}{4h}\right)^{2}}{\left(1 + \frac{\alpha}{4h}\right)^{2}} dt^{2} - \left(1 + \frac{\alpha}{4h}\right)^{4} \left[dh^{2} + h^{2} \left(d\theta^{2} + \sin^{2}\theta d\varphi^{2}\right)\right]$$

$$h = h(r) \left[\left|r - r_{o}\right|^{n} + \left(\frac{\alpha}{4}\right)^{n}\right]^{\frac{1}{n}} \qquad R_{c}(r) = h(r) \left(1 + \frac{\alpha}{4h(r)}\right)^{2} = \frac{1}{\sqrt{K(r)}}$$

$$R_{p}(r) = h(r) + \frac{\alpha}{2} \ln\left(\frac{4h(r)}{\alpha}\right) - \frac{\alpha^{2}}{8h(r)} + \frac{\alpha}{4}$$

$$R_{p}(r_{o}) = 0, \quad R_{c}(r_{o}) = \alpha, \quad K(r_{o}) = \alpha^{-2} \quad \forall r_{o} \; \forall n$$

$$r \in \mathbf{R}, \quad n \in \mathbf{R}^+, \quad r \neq r_o$$
¹⁵/₃₅

Kerr-Newman Geometry

$$ds^{2} = \frac{\Delta}{\rho^{2}} (dt - a \sin^{2} \theta \, d\varphi^{2})^{2} - \frac{\sin^{2} \theta}{\rho^{2}} [(R_{c}^{2} + a^{2})d\varphi - a \, dt]^{2} - \frac{\rho^{2}}{\Delta} dR_{c}^{2} - \rho^{2} d\theta^{2}$$
$$R = R(r) = (|r - r_{o}|^{n} + \beta^{n})^{1/n}, \qquad \beta = \frac{\alpha}{2} + \sqrt{\frac{\alpha^{2}}{4} - (q^{2} + a^{2} \cos^{2} \theta)},$$

$$\rho^{2} = R_{c}^{2} + a^{2} \cos^{2} \theta, \qquad \Delta = R_{c}^{2} - \alpha R_{c} + q^{2} + a^{2},$$

$$a=\frac{2L}{\alpha}, \qquad q^2+a^2<\frac{\alpha^2}{4},$$

$$r \in \mathbf{R}, n \in \mathbf{R}^+, r \neq r_o$$

n and r_o arbitrary constants; $R(r_o) = \beta$ is a scalar invariant.

There is no black hole.

Newtonian Approximation

$$ds^{2} = \left(c^{2} - \frac{2m}{r}\right)dt^{2} - \left(1 - \frac{2m}{rc^{2}}\right)^{-1}dr^{2} - r^{2}\left(d\theta^{2} + \sin\theta^{2}d\varphi^{2}\right)$$

Claimed to be a one-body configuration.

But escape velocity

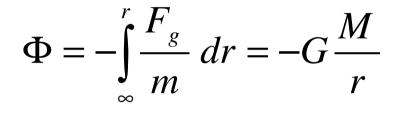
$$v = \sqrt{\frac{2Gm}{r}}$$

is a two-body concept.

Newton's Gravitation

$$F_g = -G\frac{mM}{r^2}$$

is a two-body theory.

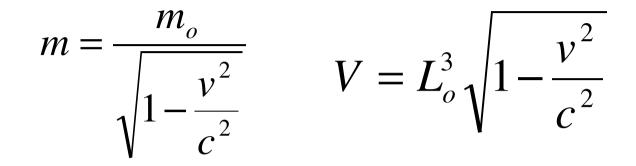


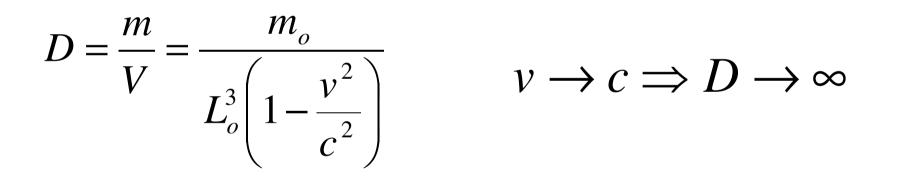
Newtonian potential is a two-body concept.

$$P = m\Phi = -G\frac{mM}{r}$$

Newton's potential energy involves two bodies.

Special Relativity

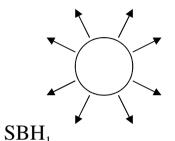




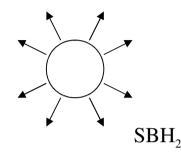
Infinite density is forbidden.

Black Hole Interactions

 $SBH \equiv Schwarzschild Black$ Hole



Ric = 0: no matter present out here



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SBH₁ is in the **matter-less** spacetime of SBH₂ and SBH₂ is in the **matter-less** spacetime of SBH₁. Thus the two black holes mutually persist in and mutually interact in a mutual spacetime that **by construction contains no matter**! Contra hyp!

 $Ric = R_{\mu\nu} = 0$ is <u>**not**</u> a two-body configuration.

The Principle of Superposition does not apply.

Riemann Tensor Scalar Curvature Invariant (Kretschmann Scalar)

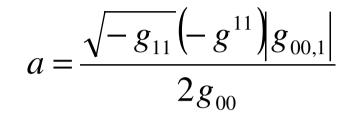
$$f = R_{\alpha\beta\gamma\delta} R^{\alpha\beta\gamma\delta}$$

$$f = 12\alpha^{2}K^{3} = \frac{12\alpha^{2}}{R_{c}^{6}} = \frac{12\alpha^{2}}{\left(\left|r - r_{o}\right|^{n} + \alpha^{n}\right)^{6/n}}$$

$$f(r_o) = \frac{12}{\alpha^4} \quad \forall r_o \; \forall n$$

 $R_p(r_o) = 0, \quad R_c(r_o) = \alpha, \quad K(r_o) = \alpha^{-2}$

Radial Geodesic Acceleration of a Point



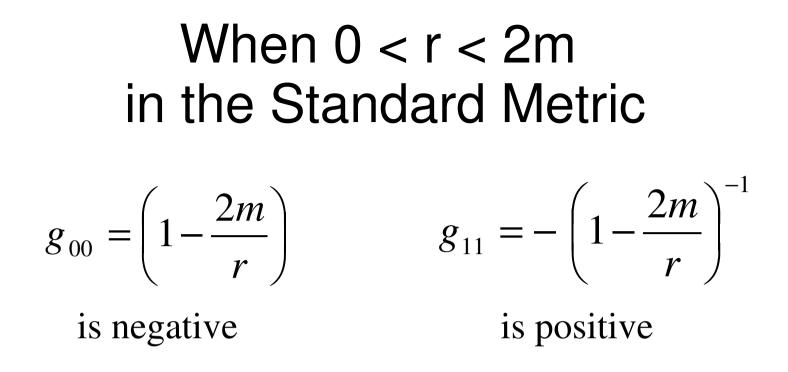
Doughty's General Expression

$$a = \frac{\alpha}{R_c^{\frac{3}{2}}(r)\sqrt{R_c(r)-\alpha}}$$

$$\lim_{r \to r_o^{\pm}} R_c(r) = \alpha$$
$$r \to r_o^{\pm} \Longrightarrow a \to \infty \quad \forall r_o \ \forall n$$

$$a = \frac{2m}{r^{\frac{3}{2}}\sqrt{r-2m}}$$

At r = 2m the acceleration is infinite, where it is claimed by the astrophysical scientists that <u>there is no matter</u>!



The signature changes from (+,-,-,-) to (-,+,-,-);

so the rôles of t and r are **interchanged**.

To amplify this set $t = r^*$, $r = t^*$.

Non-Static Solution to a Static Problem

$$ds^{2} = \left(1 - \frac{2m}{t^{*}}\right) dr^{*2} - \left(1 - \frac{2m}{t^{*}}\right)^{-1} dt^{*2} - t^{*2} \left(d\theta^{2} + \sin^{2}\theta d\varphi^{2}\right)$$

 $0 < t^* < 2m$

The signature is (-,+,-,-).

This metric is a function of the timelike *t**: a time-dependent metric. **Contra hyp**!

Principle of Equivalence

According to Einstein, his 'Principle of Equivalence' and his laws of Special Relativity **must manifest** in a sufficiently small region of his gravitational field.

Both the 'Principle of Equivalence' and the laws of Special Relativity are defined in terms of **multiple** arbitrarily large finite masses.

But Ric = 0 is a spacetime that by construction **contains no matter**! Thus Ric = 0 violates the physical basis of Einstein's gravitational field.

Empty Spacetimes $R_{\mu\nu} = \lambda g_{\mu\nu}$

- λ *is* the 'cosmological constant'.
- The spacetime contains no matter by construction $(T_{\mu\nu} = 0)$.

• Is satisfied by de Sitter's **empty** universe (which can be obtained by setting m = 0 in the so-called 'Schwarzschild/de Sitter metric'.

$$R_{\mu\nu}=0$$

• The spacetime contains no matter by construction $(T_{\mu\nu} = 0)$.

• Alleged to contain matter (*post hoc*), denoted by *m* in the so-called "Schwarzschild solution".

Thus $T_{\mu\nu} = 0$ is alleged to both support material cause and to preclude material cause. ²⁶/₃₅

Empty Spacetimes

$$ds^{2} = \left(1 - \frac{2m}{r} - \frac{\lambda}{3}r^{2}\right)dt^{2} - \left(1 - \frac{2m}{r} - \frac{\lambda}{3}r^{2}\right)^{-1}dr^{2} - r^{2}\left(d\theta^{2} + \sin^{2}\theta d\varphi^{2}\right)$$

The "Schwarzschild – de Sitter" metric: $T_{\mu\nu} = 0$,

but said to contain material cause, m.

$$ds^{2} = \left(1 - \frac{\lambda}{3}r^{2}\right)dt^{2} - \left(1 - \frac{\lambda}{3}r^{2}\right)^{-1}dr^{2} - r^{2}\left(d\theta^{2} + \sin^{2}\theta d\varphi^{2}\right)$$

de Sitter's completely empty world: $T_{\mu\nu} = 0$,

said to contain no material cause.

Einstein's Field Equations

$$\frac{G_{\mu\nu}}{\kappa} + T_{\mu\nu} = 0$$

• $G_{\mu\nu}/\kappa$ are the components of a gravitational energy tensor.

- •The total energy of the gravitational field is **always 0**.
- •The $G_{\mu\nu}/\kappa$ and the $T_{\mu\nu}$ must vanish identically.
- •Gravitational energy cannot be localised.
- •Einstein's gravitational field **violates** the usual conservation of energy and momentum established by experiment.

Einstein's pseudo-tensor

$$\sqrt{-g}t_{\nu}^{\mu} = \frac{1}{2} \left(\delta_{\nu}^{\mu}L - \frac{\partial L}{\partial g_{,\mu}^{\sigma\rho}} g_{,\nu}^{\sigma\rho} \right)$$

$$L = -g^{\alpha\beta} \left(\Gamma^{\gamma}_{\alpha\kappa} \Gamma^{\kappa}_{\beta\gamma} - \Gamma^{\gamma}_{\alpha\beta} \Gamma^{\kappa}_{\lambda\kappa} \right)$$

definition of Einstein's pseudo-tensor

Einstein's pseudo-tensor is meaningless

$$\sqrt{-g} t^{\mu}_{\mu} = \frac{1}{2} \left(4L - \frac{\partial L}{\partial g^{\sigma\rho}_{,\mu}} g^{\sigma\rho}_{,\mu} \right)$$

$$\frac{\partial L}{\partial g_{,\mu}^{\,\sigma\rho}}g_{,\mu}^{\,\sigma\rho} = 2L$$

So the invariant is L.

But such invariants do not exist!

Linearisation and Perturbation

Linearisation of Einstein's field equations is inadmissible because linearisation implies the existence of a tensor which, except for the trivial case of being precisely zero, **does not exist**!

(proven by the German mathematician Hermann Weyl, in 1944)

Conclusions

- "Schwarzschild's solution" is not Schwarzschild's solution. Black holes are not predicated by Schwarzschild's solution.
- The quantity r in "Schwarzschild's solution" is the inverse square root of the Gaussian curvature of a spherically symmetric geodesic surface in the spatial section and is therefore not itself a distance of any kind in the manifold.
- *Ric* = 0 is a spacetime that by construction contains no matter and therefore **violates** the physical principles of General Relativity (the 'Principle of Equivalence' and the laws of Special Relativity **cannot manifest** in a spacetime that **by construction contains no matter**).

Conclusions

•General Relativity **does not** predict black holes. **Black holes do not exist**.

•Einstein's gravitational field **violates** the usual conservation of energy and momentum established by experiment. **Einstein** gravitational waves do not exist.

•Despite the numerous claims of the astrophysical scientists, no black holes and no Einstein gravitational waves have been detected. Gravity Probe B **did not detect** the Lense-Thirring effect. The international efforts to detect black holes and Einstein gravitational waves **are destined to detect nothing**.

Acknowledgements

I thank **Monika Vandory** for providing the financial means for me to attend this conference. Without her generous benefaction I would not have been able to be here.

Dedication

For my late brother,

Paul Raymond Crothers

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